

# Development of a new Solar Radiance-To-Flux Conversion to Improve SW Flux Estimations

–final presentation–

Florian Tornow<sup>1,2,3</sup>, Nils Madenach<sup>1</sup>, René Preusker<sup>1</sup>, Howard Barker<sup>4</sup>, Jason Cole<sup>4</sup>, Almudena Velázquez Blázquez<sup>5</sup>, Carlos Domenech<sup>6</sup> and Jürgen Fischer<sup>1</sup>

<sup>1</sup>Freie Universität Berlin

<sup>2</sup>Columbia University, New York

<sup>3</sup>NASA Goddard Institute for Space Studies, New York

<sup>4</sup>Environment and Climate Change Canada, Toronto

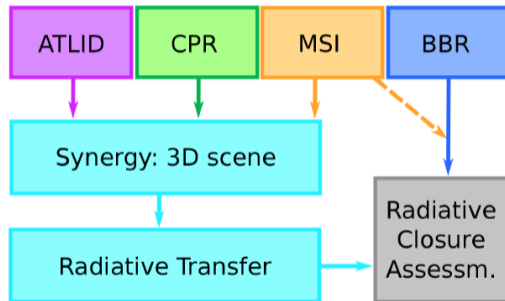
<sup>5</sup>Royal Meteorological Institute, Brussels

<sup>6</sup>GMV, Madrid

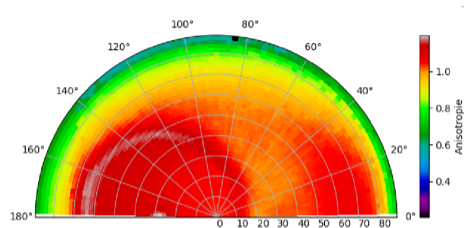
*nils.madenach@wew.fu-berlin.de*

March 10, 2021

- | EarhtCare's **Radiative closure** successful if  $F = F_{RT} - F_{BBR}$  within  $10 \text{ Wm}^2$

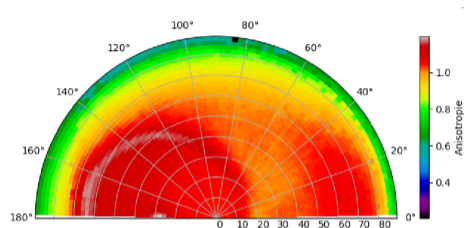


- | EarhtCare's **radiative closure** successful if  $F = F_{RT} - F_{BBR}$  within  $10 \text{ Wm}^2$
- | **BBR Radiance-to- flux conversion** should be as good as possible



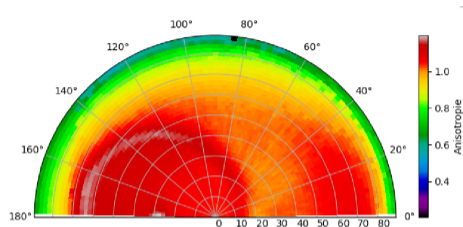
SW radiance angular dependency of an overcast scene with  $\tau = 12$

- | EarhtCare's **radiative closure** successful if  $F = F_{RT} - F_{BBR}$  within  $10 \text{ Wm}^2$
- | **BBR Radiance-to- flux conversion** should be **as good as possible**
- | Quality of BBR SW flux estimates depends on underlying **angular distribution model (ADM)**



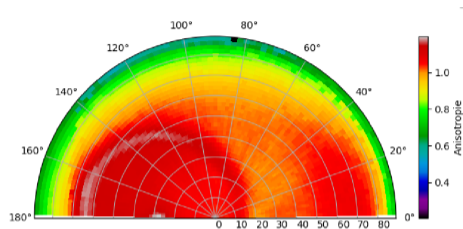
SW radiance angular dependency of an overcast scene with  $\tau = 12$

- | EarhtCare's **radiative closure** successful if  $F = F_{RT} - F_{BBR}$  within  $10 \text{ Wm}^2$
- | **BBR Radiance-to- flux conversion** should be **as good as possible**
- | Quality of BBR SW flux estimates depends on underlying **angular distribution model (ADM)**
  - | ADMs account for the **angular dependency** of radiation field (anisotropy)



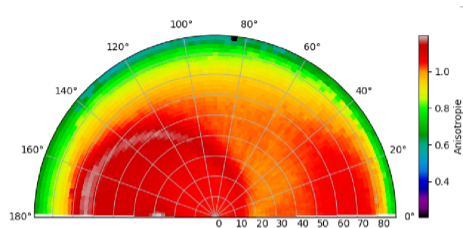
SW radiance angular dependency of an overcast scene with  $\tau = 12$

- | EarhtCare's **radiative closure** successful if  $F = F_{RT} - F_{BBR}$  within  $10 \text{ Wm}^2$
- | **BBR Radiance-to- flux conversion** should be **as good as possible**
- | Quality of BBR SW flux estimates depends on underlying **angular distribution model (ADM)**
  - | ADMs account for the **angular dependency** of radiation field (anisotropy)
  - | **State-of-art** sigmoidal approach is a function of  $\mu$ ,  $f$  and  $w_{10m}$



SW radiance angular dependency of an overcast scene with  $\tau = 12$

- | EarhtCare's **radiative closure** successful if  $F = F_{RT} - F_{BBR}$  within  $10 \text{ Wm}^2$
- | **BBR Radiance-to- flux conversion** should be **as good as possible**
- | Quality of BBR SW flux estimates depends on underlying **angular distribution model (ADM)**
  - | ADMs account for the **angular dependency** of radiation field (anisotropy)
  - | **State-of-art** sigmoidal approach is a function of  $\mu$ ,  $f$  and  $w_{10m}$
  - | **lack** of important **dependencies** of e.g. **cloud micro-physics**



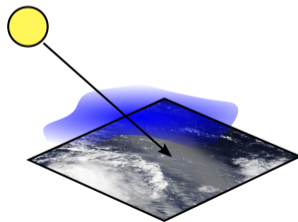
SW radiance angular dependency of an overcast scene with  $\tau = 12$

- Incorporated **additional dependencies** of SW radiance to cloud micro-physics (via  $R_{eff}$ )

$$\log I(s_i, v_i) = \log S_0 + \log \tau + 2 \log CTWV$$

$S_0$  = solar constant  
 $\tau$  = footprint albedo

$s_i, v_i$  = sun-observer angles

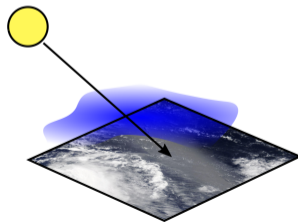




- | Incorporated **additional dependencies** of SW radiance to cloud micro-physics (via  $R_{eff}$ )
- | and  $CTWV$

$$\log I(s_i, v_i) = \log S_0 + \log \tau + \log 2 CTWV$$

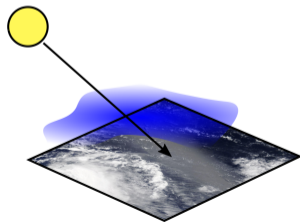
$S_0$  = solar constant  
 $\tau$  = footprint albedo  
 $s_i, v_i$  = sun-observer angles



- | Incorporated **additional dependencies** of SW radiance to cloud micro-physics (via  $R_{eff}$ )
- | and  $CTWV$
- | Relates quantities **linearly** and more continuously to **measured radiances**

$$\log I(s_i, v_i) = \log S_0 + \log \tau + 2 CTWV$$

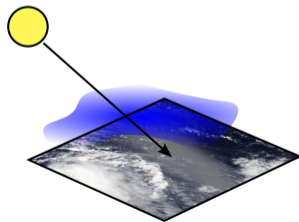
$S_0$  = solar constant  
 $\tau$  = footprint albedo  
 $s_i, v_i$  = sun-observer angles

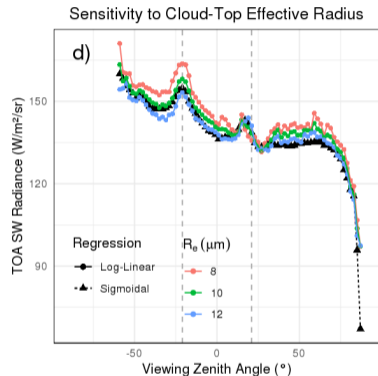
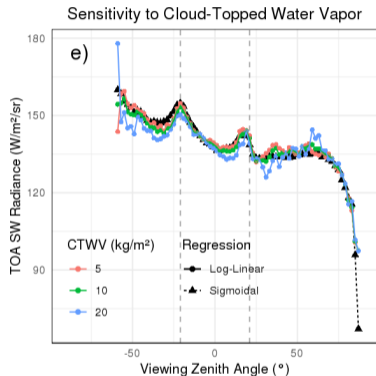


- | Incorporated **additional dependencies** of SW radiance to cloud micro-physics (via  $R_{eff}$ ) and  $CTWV$
- | Relates quantities **linearly** and more continuously to **measured radiances**
- | **Footprint albedo** ( $f; w_{10}; ; R_{eff}$ ) is calculated using **two-stream theory**

$$\log I(s; v) = \log S_0 + \log \frac{1}{2} CTWV$$

$S_0$  = solar constant  
= footprint albedo  
 $s; v$  = sun-observer angles





$CTWV$ -sensitivity along the PP  $S = 21^\circ$ , ?

$R_{eff}$ -sensitivity along the PP  $S = 21^\circ$ , ?

- | **Instantaneous flux deviations** of up to  $25 \text{ W m}^{-2}$  when applied to CERES-MODIS and GERB-SEVIRI observations of **marine clouds**

Daily flux of log-linear minus operational approach for July 11th 2007 ?

- | **Instantaneous flux deviations** of up to  $25 \text{ W m}^{-2}$  when applied to CERES-MODIS and GERB-SEVIRI observations of **marine clouds**
- | **Deviations** associated with **extremes** in  $R_{eff}$

Daily flux of log-linear minus operational approach for July 11th 2007 ?

- | **Instantaneous flux deviations** of up to  $25 \text{ W=m}^2$  when applied to CERES-MODIS and GERB-SEVIRI observations of **marine clouds**
- | **Deviations** associated with **extremes** in  $R_{eff}$
- | Deviations can propagate to **daily means** (up to  $10 \text{ W=m}^2$ )

Daily flux of log-linear minus operational approach for July 11th 2007 ?

- | **Instantaneous flux deviations** of up to  $25 W=m^2$  when applied to CERES-MODIS and GERB-SEVIRI observations of **marine clouds**
- | **Deviations** associated with **extremes** in  $R_{eff}$
- | Deviations can propagate to **daily means** (up to  $10 W=m^2$ )
- | and **monthly means** (up to  $5 W=m^2$ )

Daily flux of log-linear minus operational approach for July 11th 2007 ?



Thank You For Your Attention!